

Tomography in the Ecliptic using HI-1

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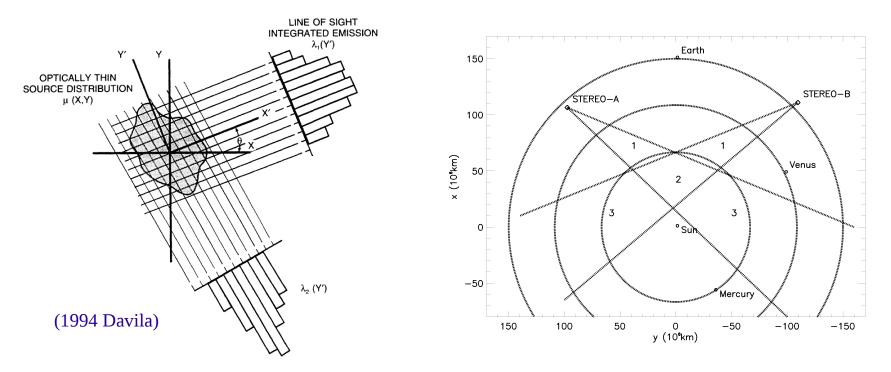
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Overview

- Theory of tomographic inversion applied to electron densities in the heliosphere
- Using HI data to formulate an inverse equation based on Thomson scattering by electrons
- Solution to the inverse equation using iterative estimation
- Separating the F- and K-corona in HI data
- Regularisation of the inverse equation (eg. positive densities, smooth gradient)
- Example of the method applied to Earth directed CMEs



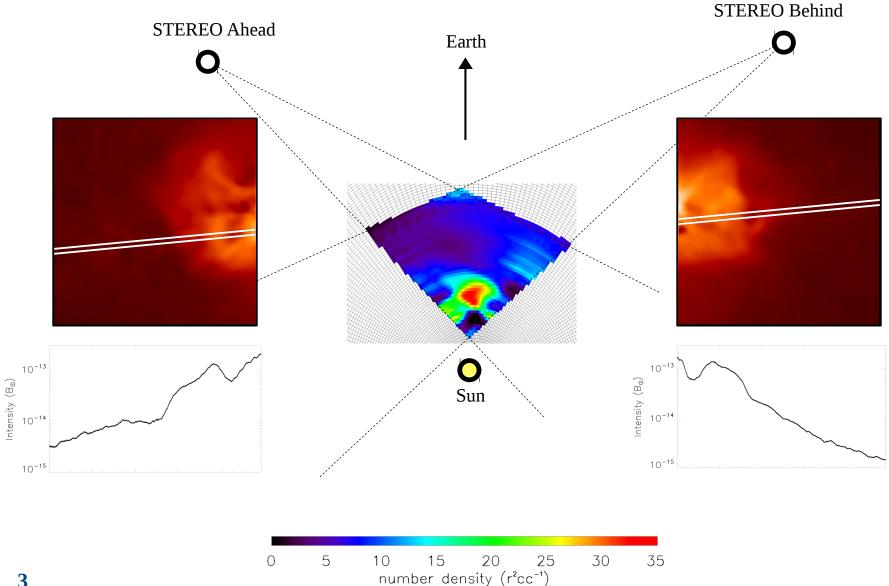
Tomography: Theory



- Single observing position produces line-of-sight ambiguities
- These may be avoided using multiple points of view
- Tomography may be performed in the region common to FOV of STEREO A and B

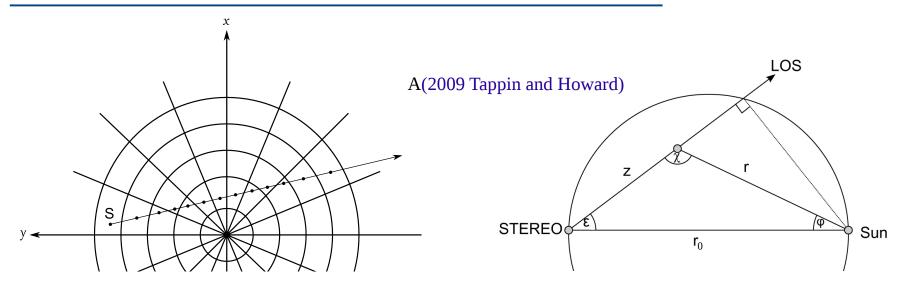


Tomography with STEREO HI



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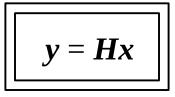
The Inverse Equation



- The equation y = Hx may be formulated by applying a grid to the heliosphere, with a resolution is 1.5° longitudinally by 1/160 AU radially
- y = data array, H = physical operator, x = density array
- $H(r, \chi)$ is based on Thomson scattering
- *H* and *y* are known, which means we can find *x*



Solving the Inverse Problem



- $y : M \sim 2000$ elements $x : N \sim 5000$ elements
- H : NxM elements... > 10⁷
- Instead we solve it using an iterative estimation method: the *Conjugate gradient algorithm*

... Too big to solve Analytically

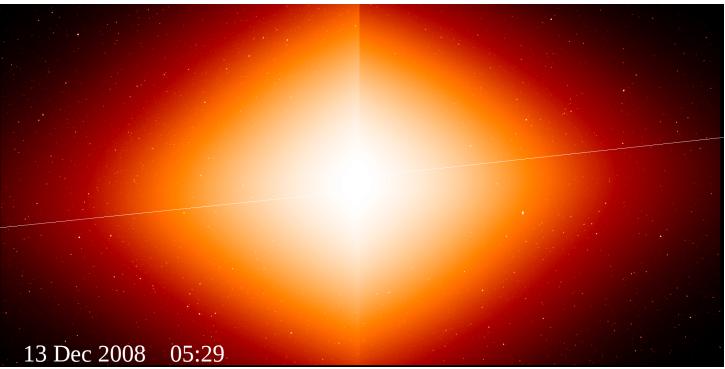
- make initial guess at density array: $x_0 = 8r^{-2}$ electrons cc⁻¹
- Find residual $r_o = y Hx_o$
- update value of x_i , such that the value of r_i is reduced
- repeat until **r**_i is sufficiently small



Background Subtraction

STEREO A

STEREO B



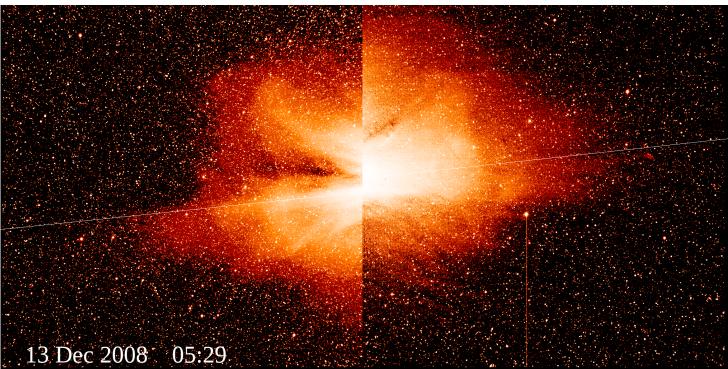
• HI-1 data dominated by F-corona (dust)



Background Subtraction

STEREO A



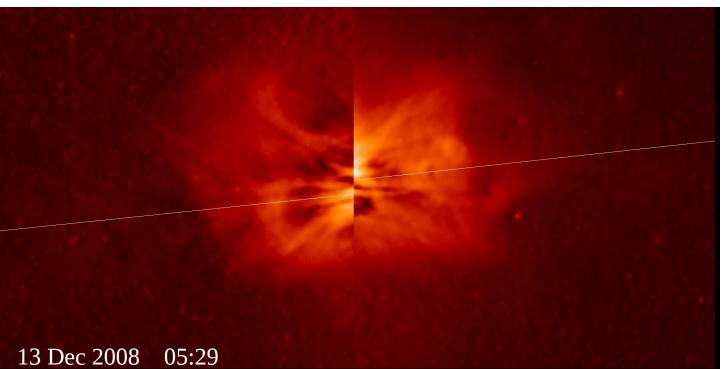


- HI-1 data dominated by F-corona (dust)
- One-day minimum subtracted from each pixel



Background Subtraction

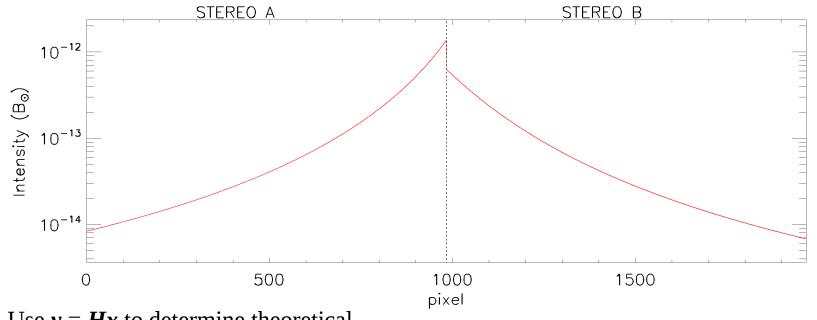
STEREO A



STEREO B

- HI-1 data dominated by F-corona (dust)
- One-day minimum subtracted from each pixel
- Apply median filter over nearby pixels

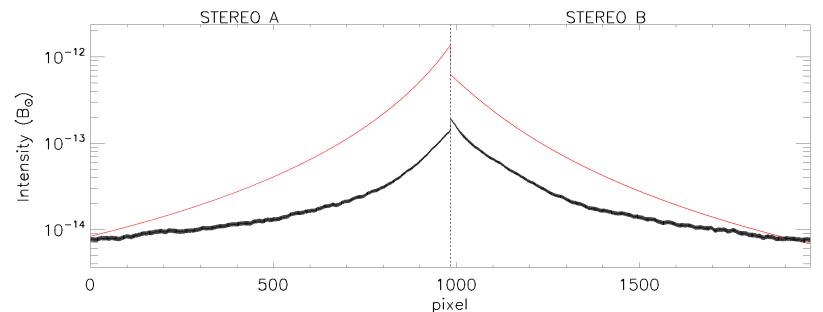




Use y = Hx to determine theoretical values of intensity

– Theoretical Intensity, $x = 8.0r^2$ electrons cc^{-1}

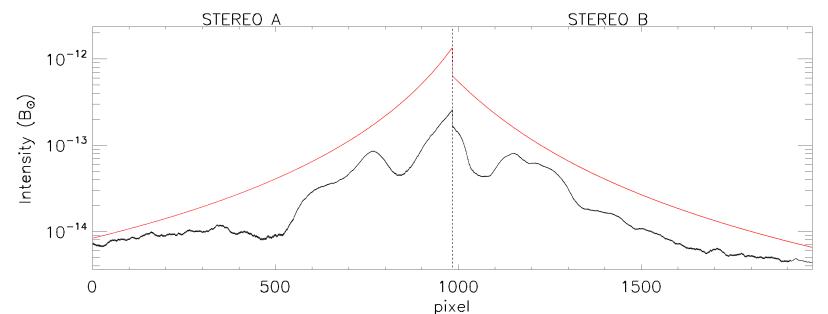




- Use *y* = *Hx* to determine theoretical values of intensity
- Ratio of *theoretical* to *observed* provides a correction factor

- Theoretical Intensity, $x = 8.0r^2$ electrons cc^{-1}
- Observed Intensity, two-week average (28 Nov – 11 Dec 2008)

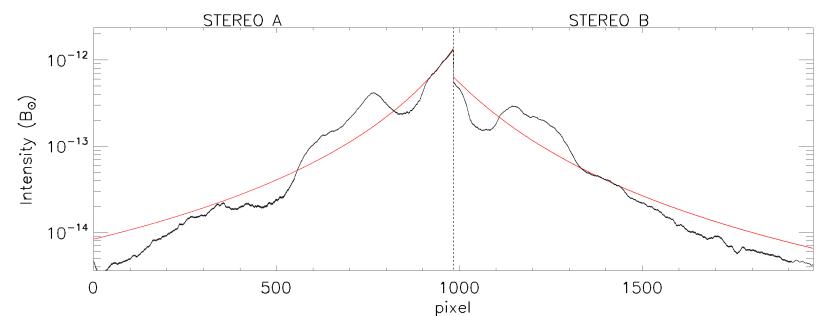




- Use *y* = *Hx* to determine theoretical values of intensity
- Ratio *theoretical* to *observed* provides a correction factor
- This may be applied to individual HI images

- Theoretical Intensity, $x = 8.0r^2$ electrons cc^{-1}
- Observed Intensity, 13 Dec 2008 05:29

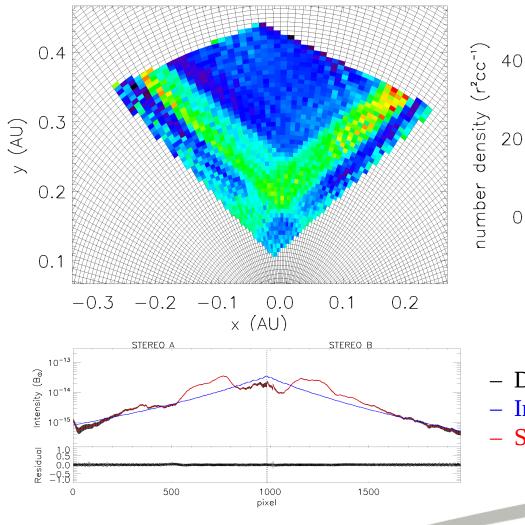




- Use *y* = *Hx* to determine theoretical values of intensity
- Ratio *theoretical* to *observed* provides a correction factor
- This may be applied to individual HI images

- Theoretical Intensity, $x = 8.0r^2$ electrons cc^{-1}
- Corrected Intensity, 13 Dec 2008 05:29

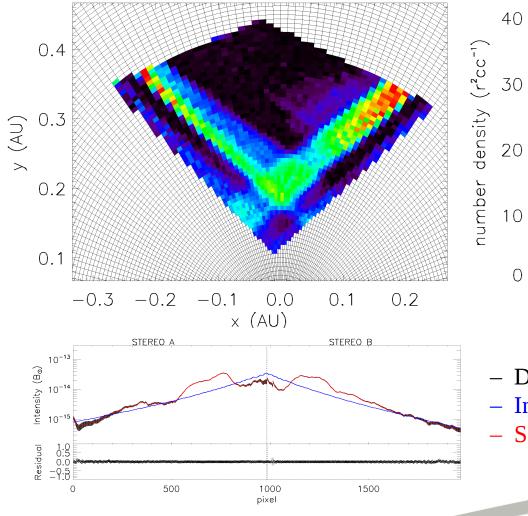




• *'Map'* represents density array, *x*

- Data
- Initial guess $x = 8.0r^2$
- Solution



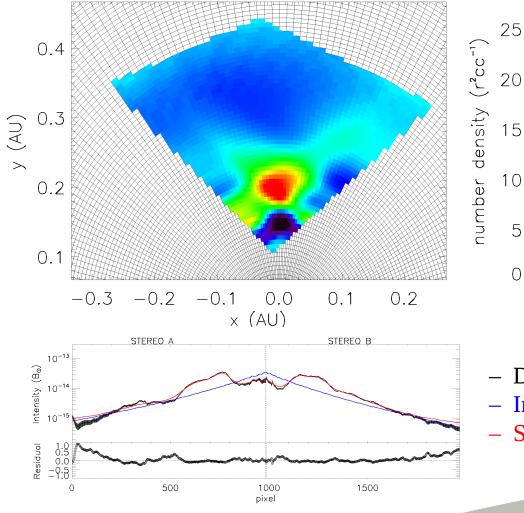


- *'Map'* represents density array, *x*
- Reset negative values to *x* after each iteration

- Data - Initial guess $x = 8.0r^2$
- Solution

 $\left(\right)$





- *'Map'* represents density array, *x*
- Reset negative values to *x* after each iteration
- Include 'regularisation' matrices, $D\varphi$ and D^2r

Data _

25

15

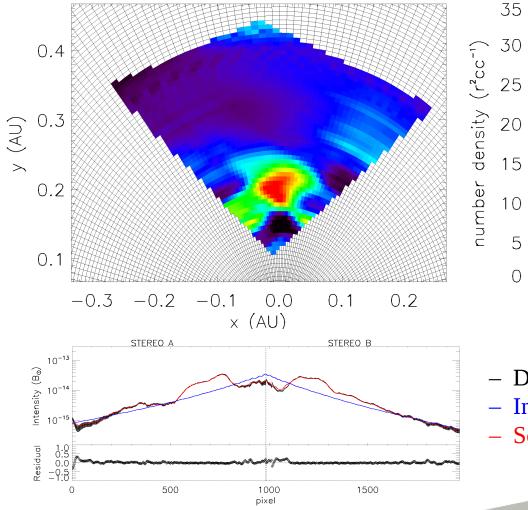
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- Initial guess $x = 8.0r^2$
- Solution





- '*Map*' represents density array, *x*
- Reset negative values to x_{a} after each iteration
- Include 'regularisation' matrices, $D\varphi$ and D^2r
- Add weighting each pixel to account for range of intensities (factor $\sim 10^2$)

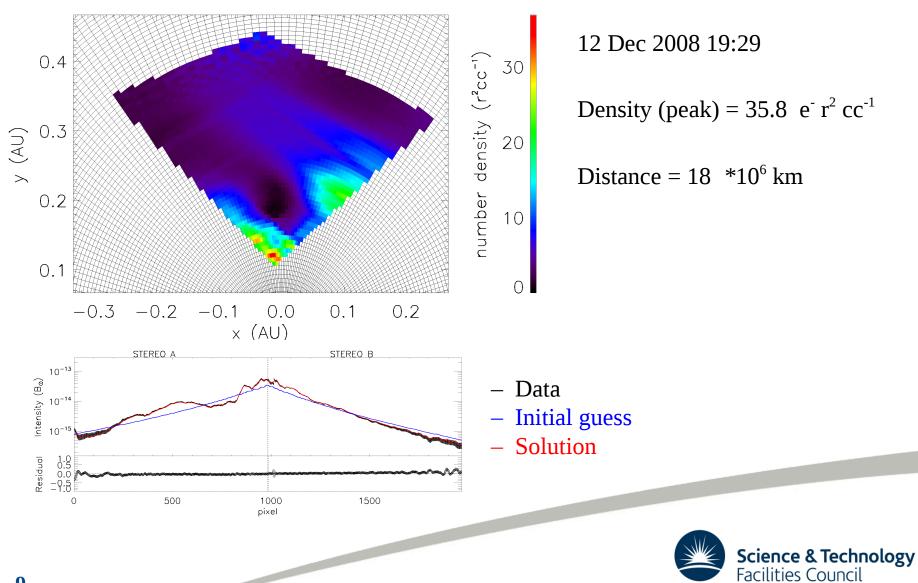
Data _

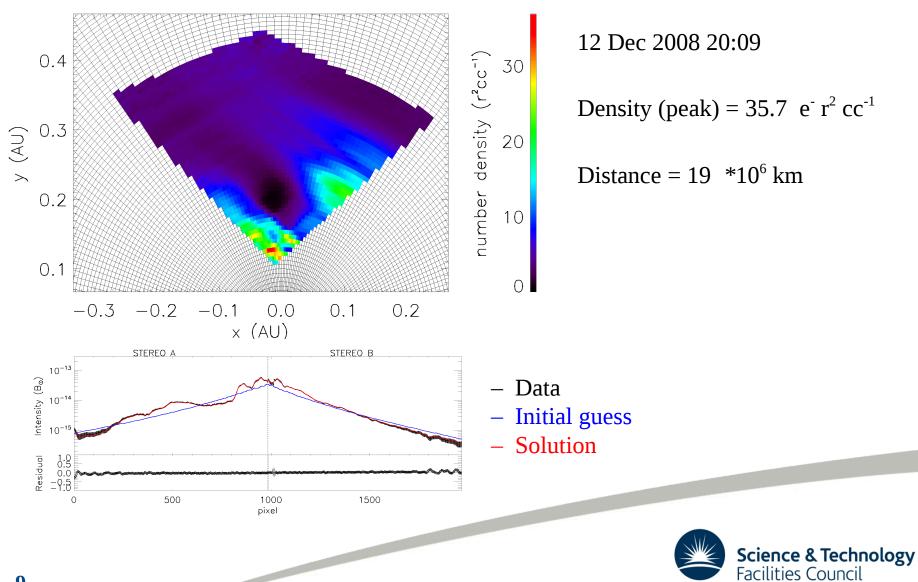
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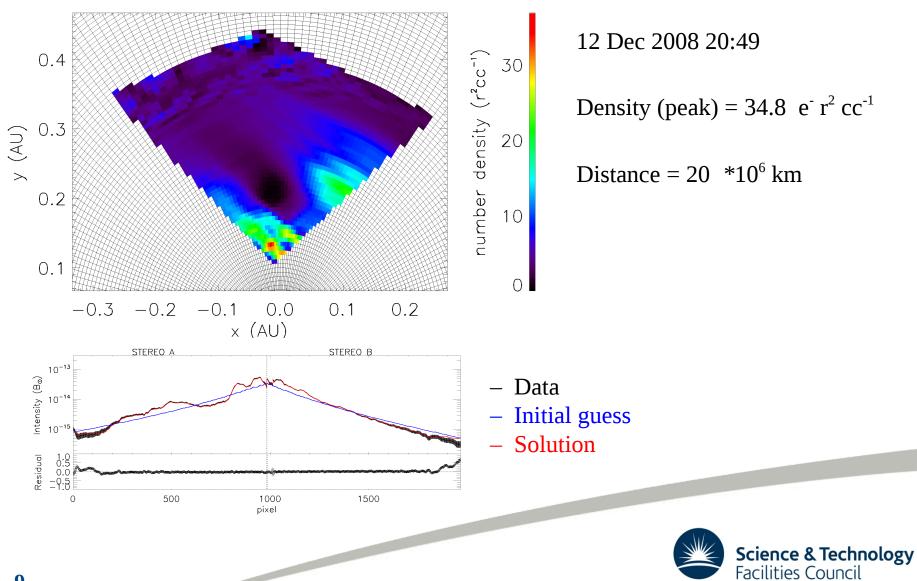
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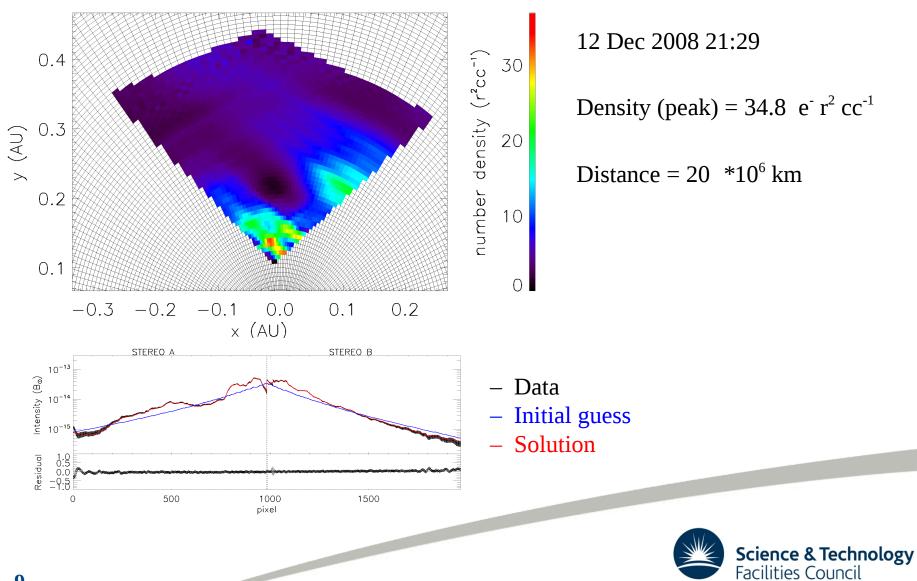
- Initial guess $x = 8.0r^2$
- Solution

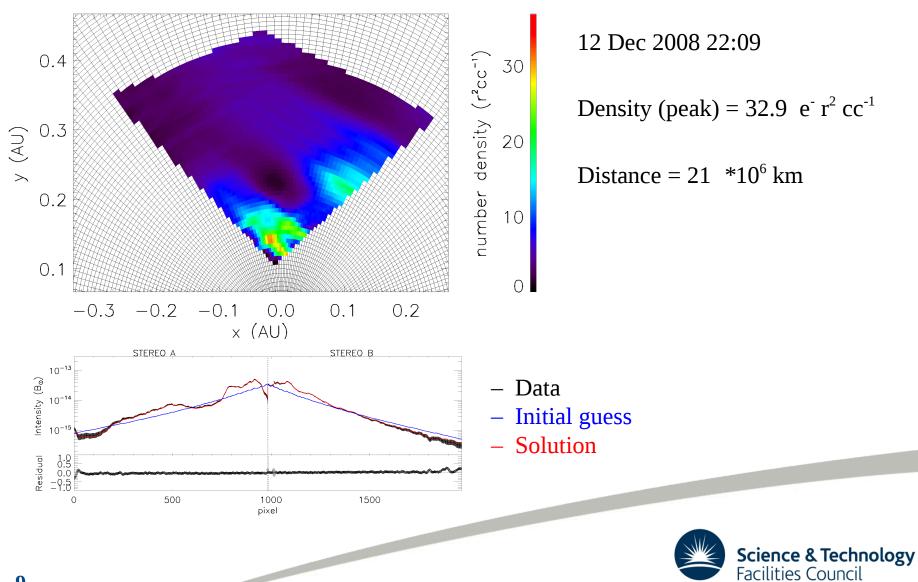


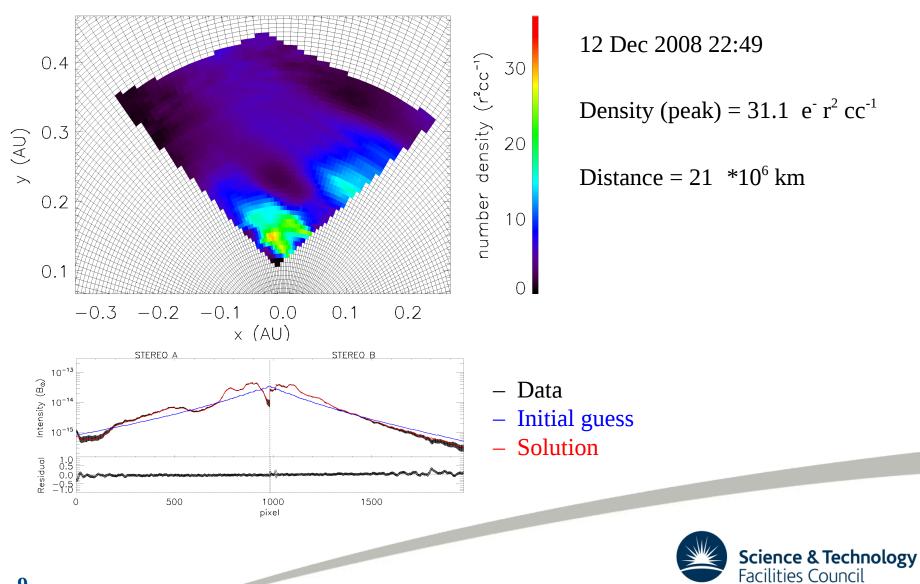


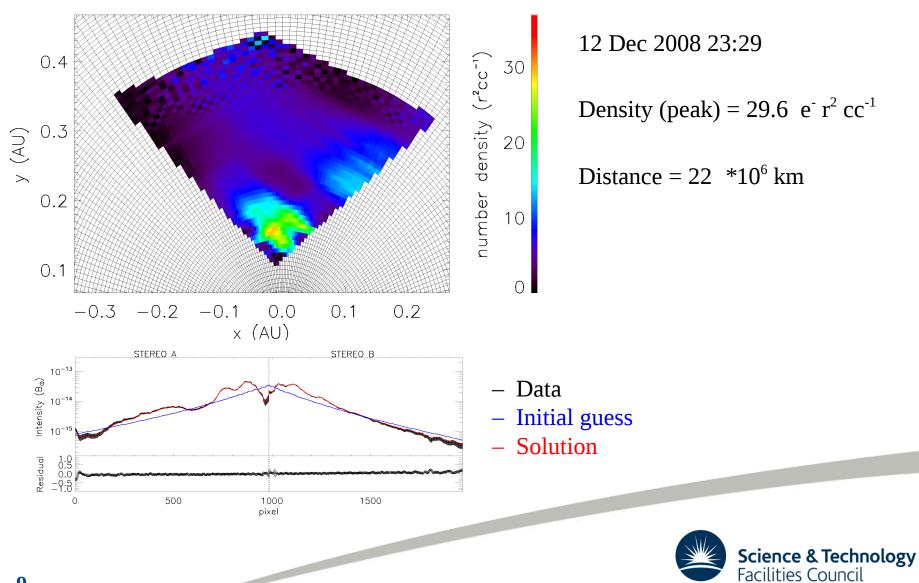


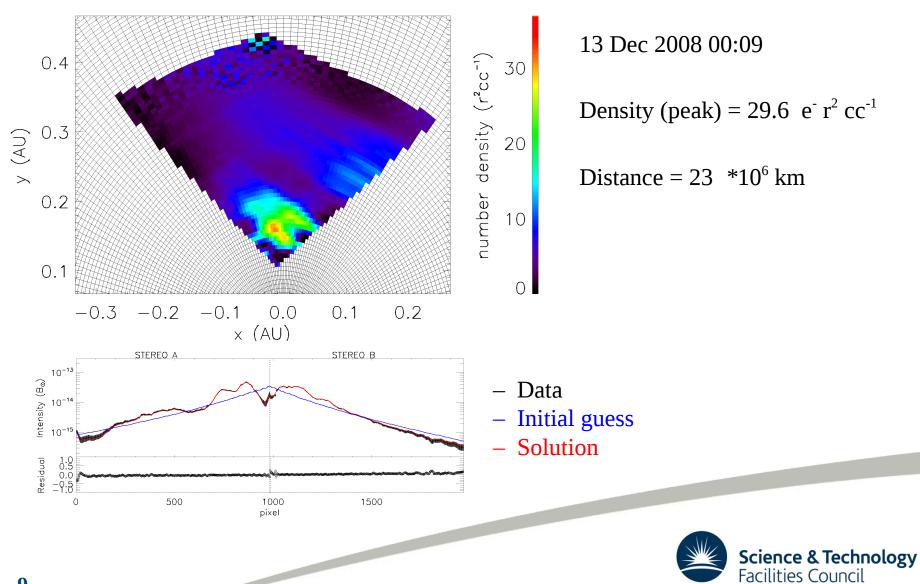


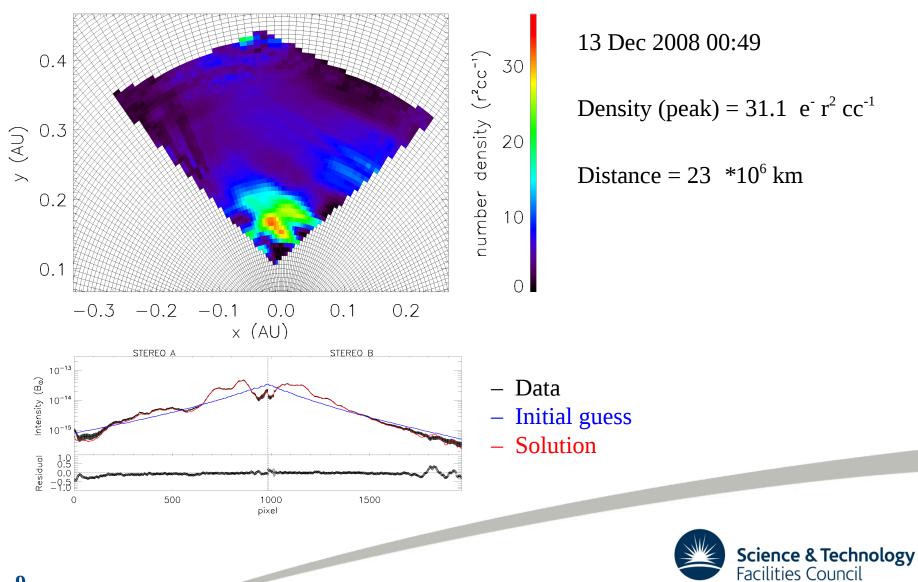


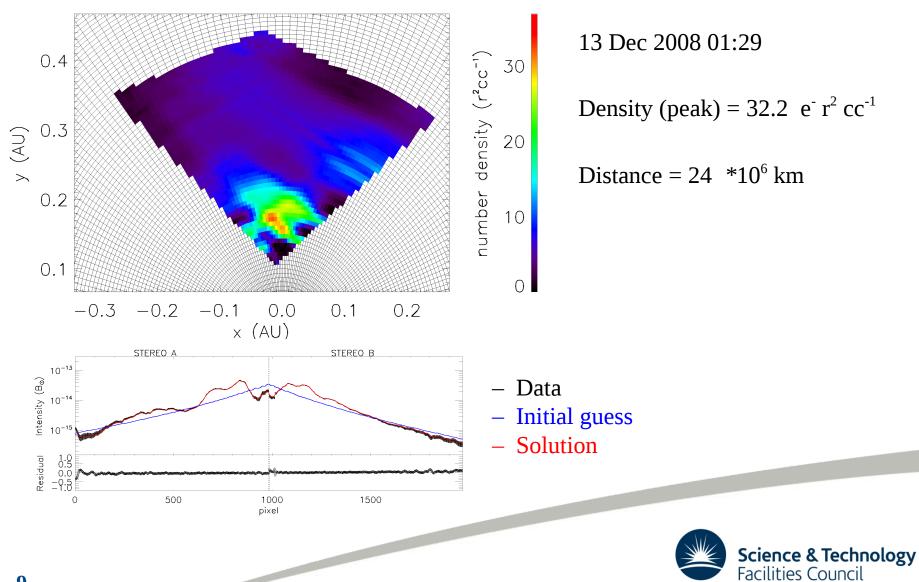


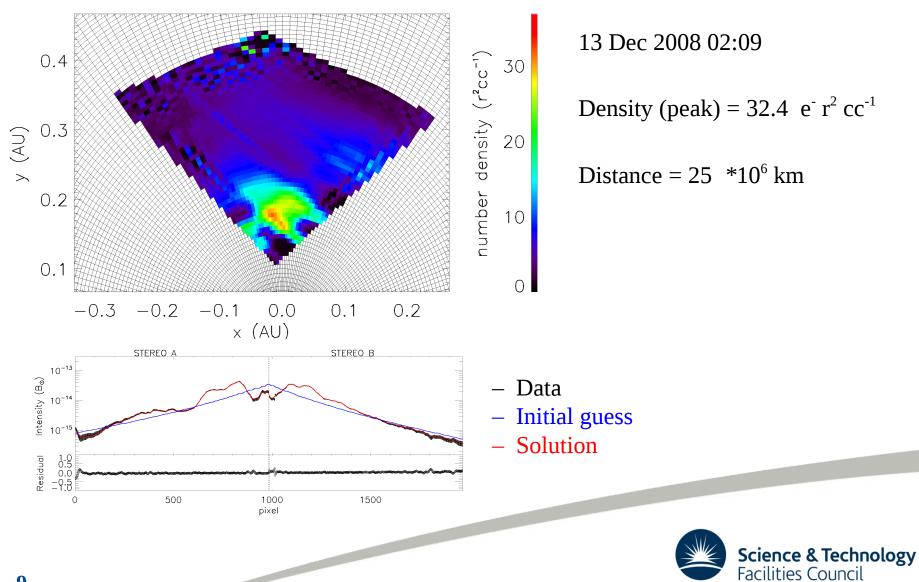


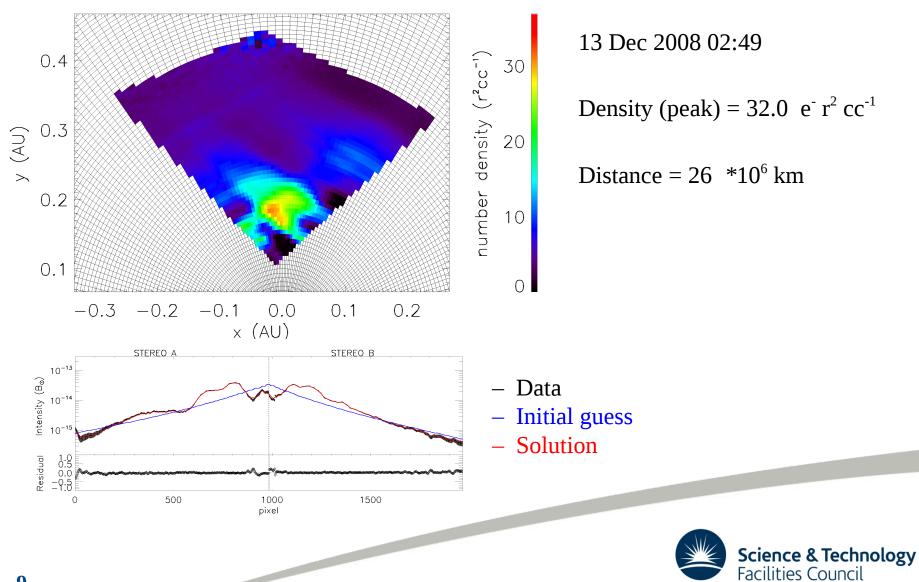


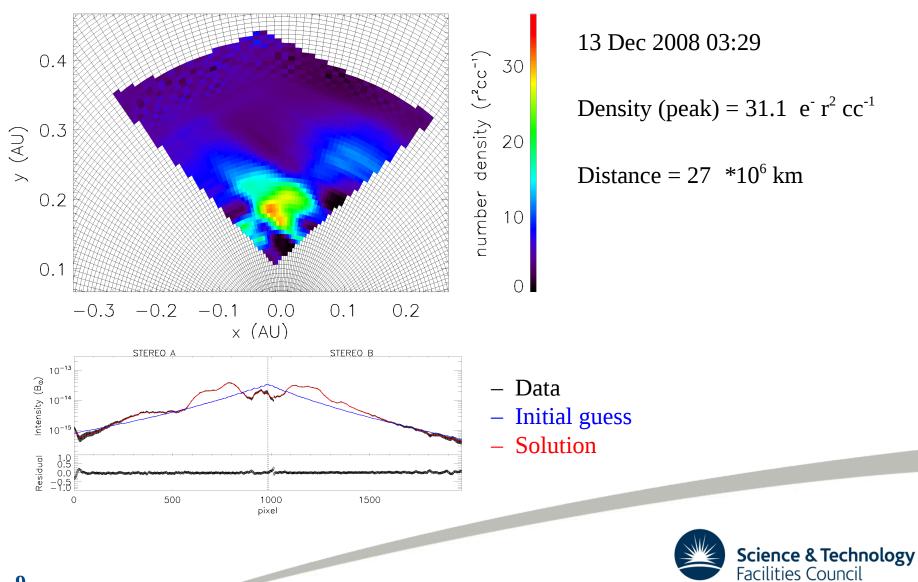


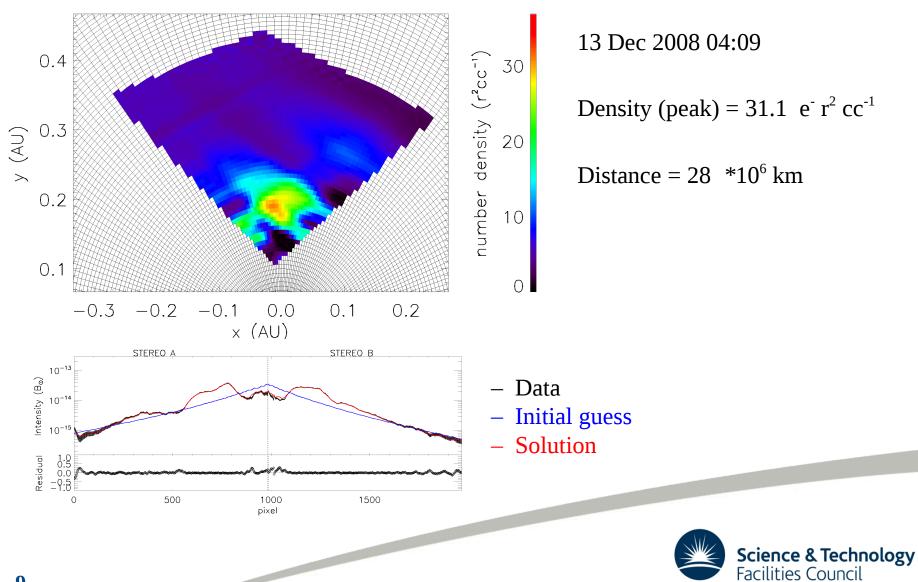


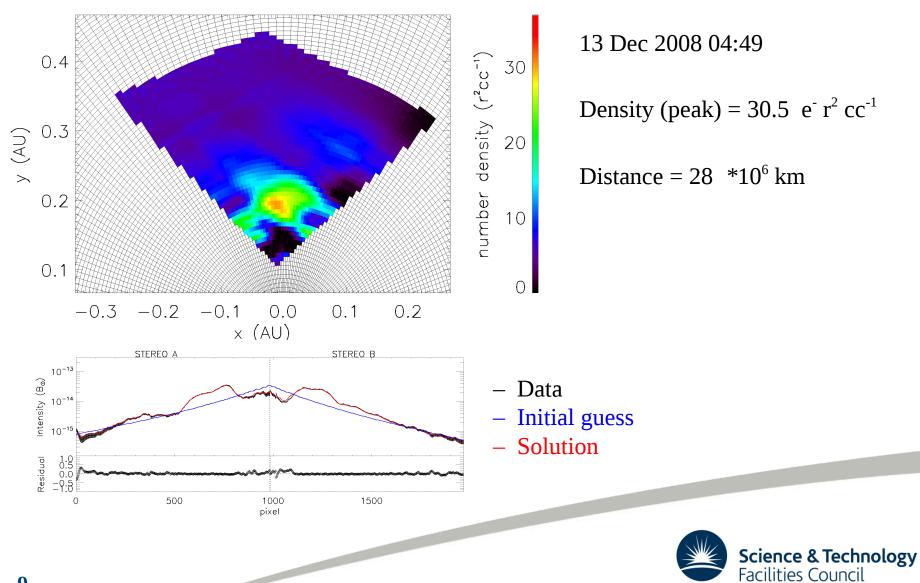


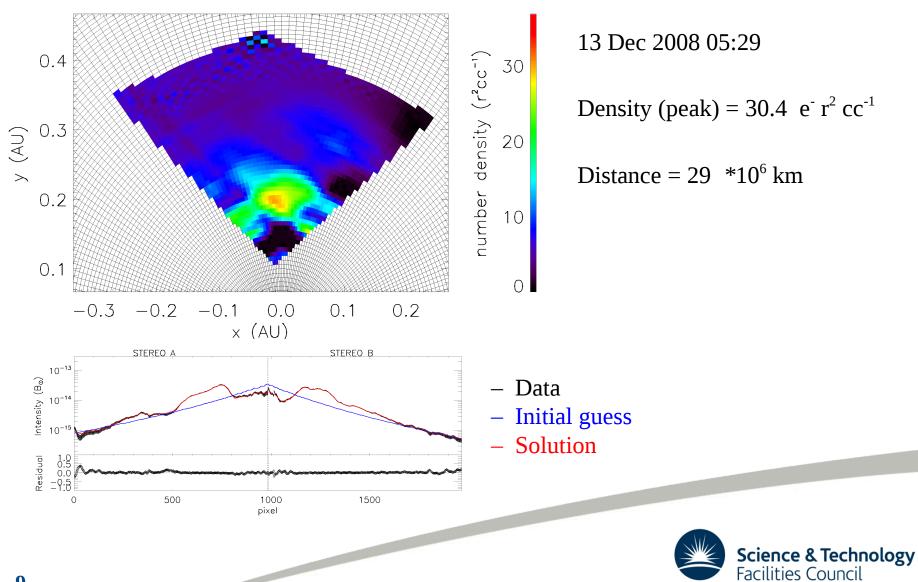


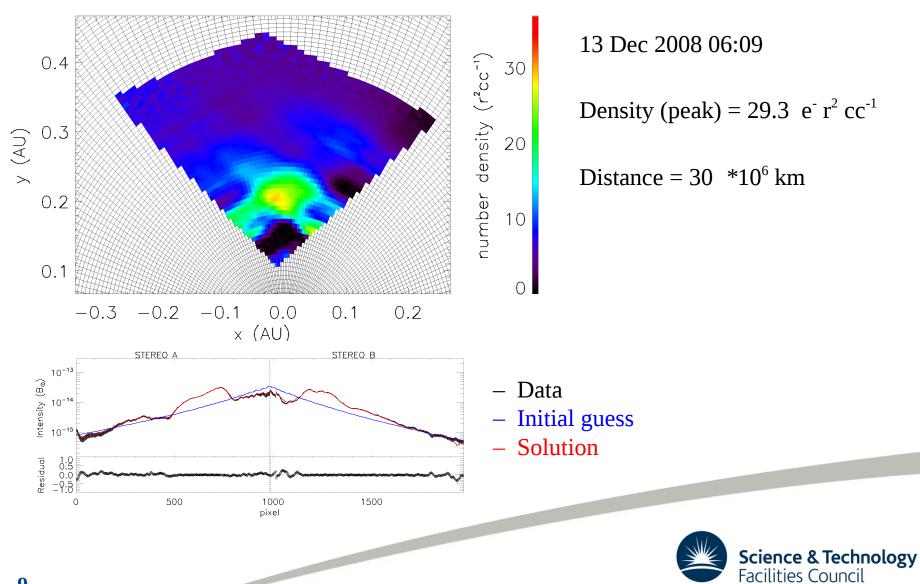


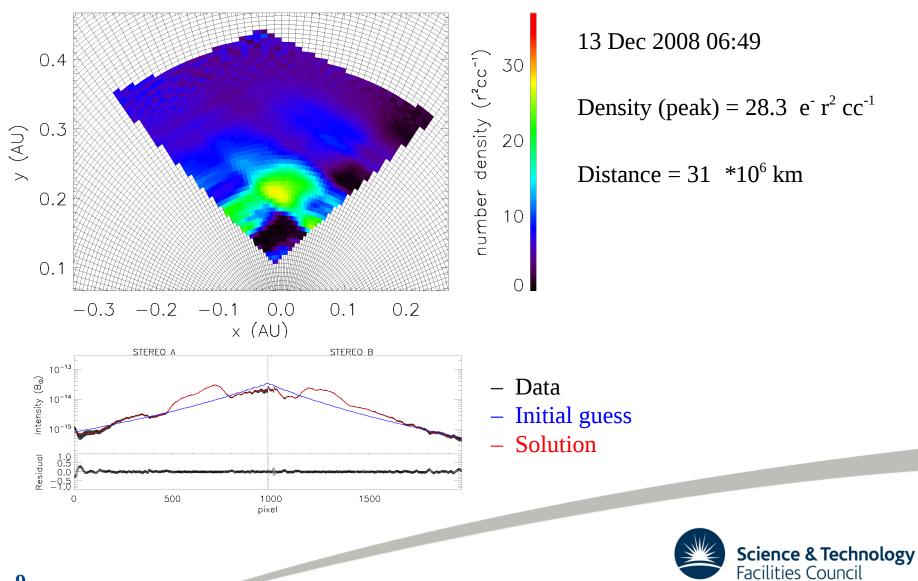


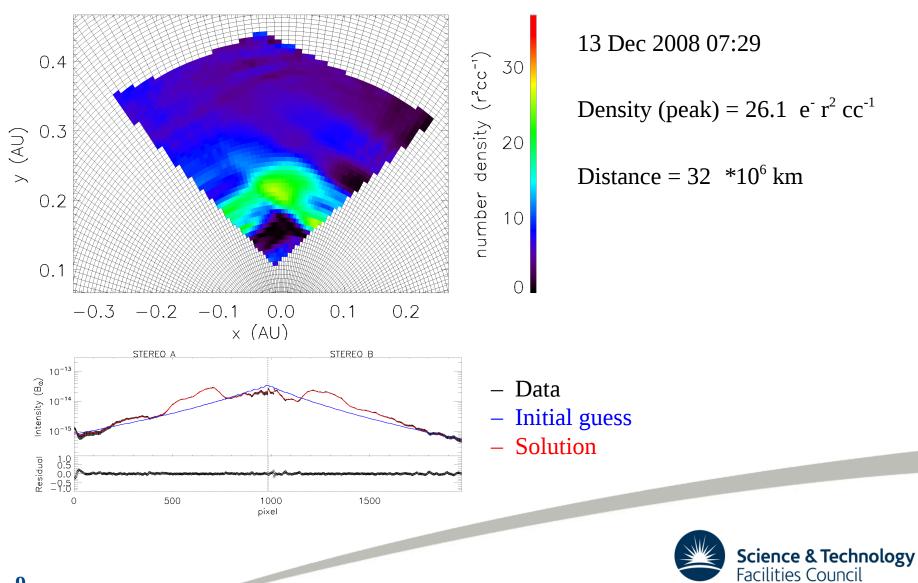


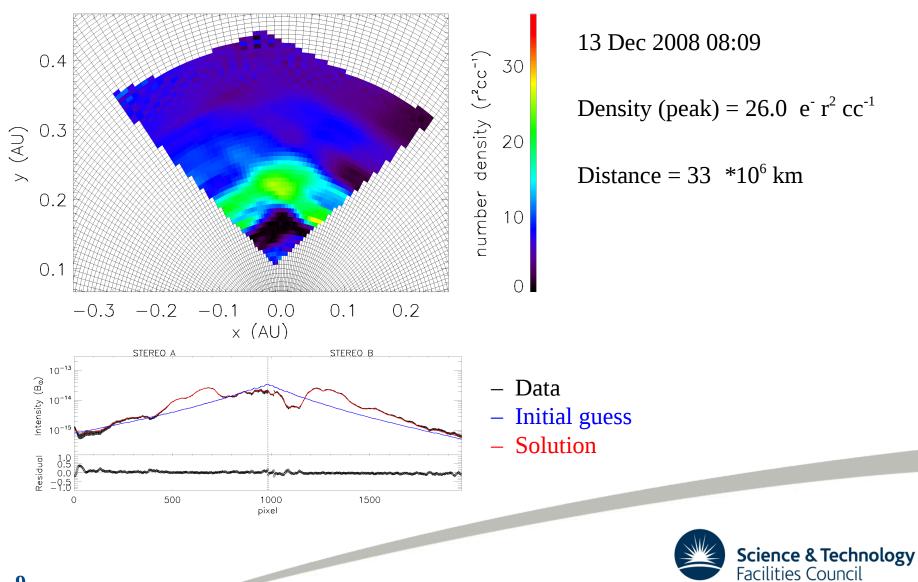


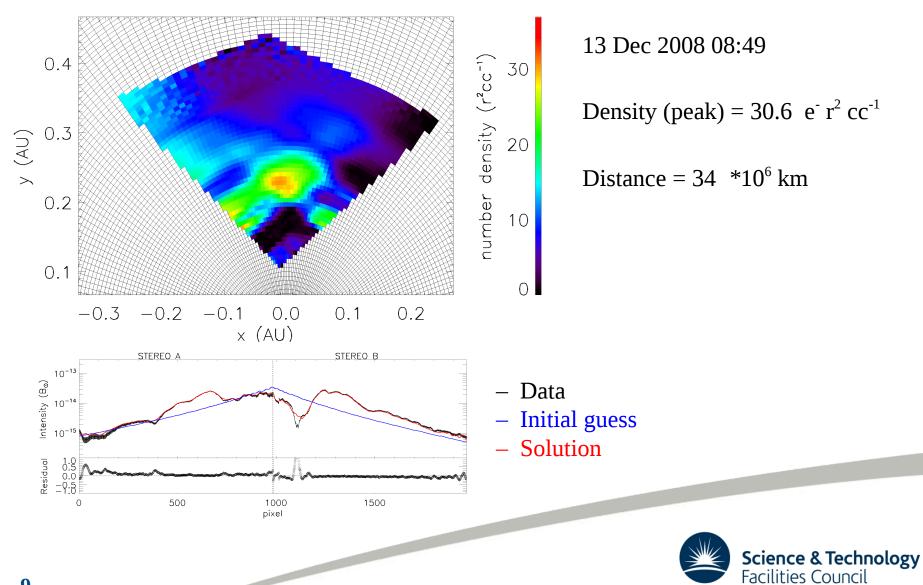


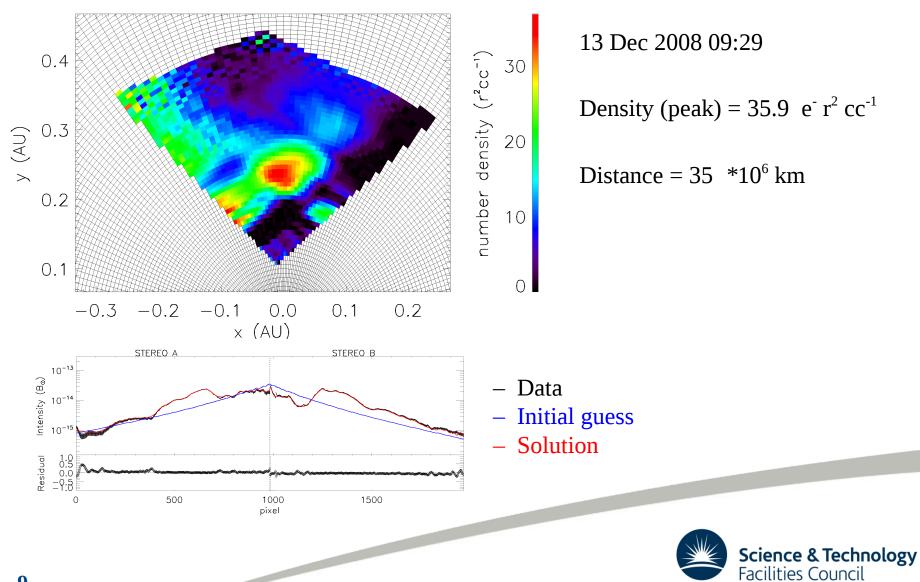


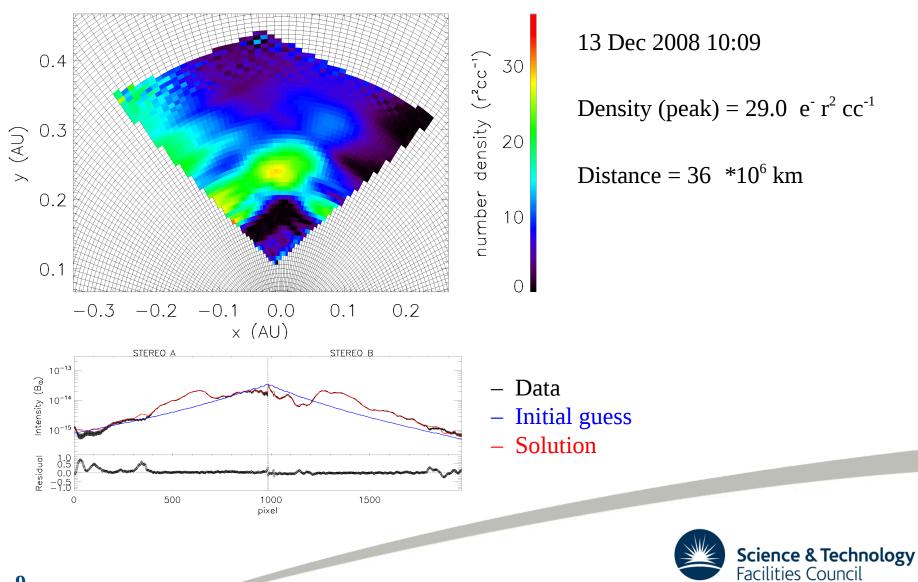


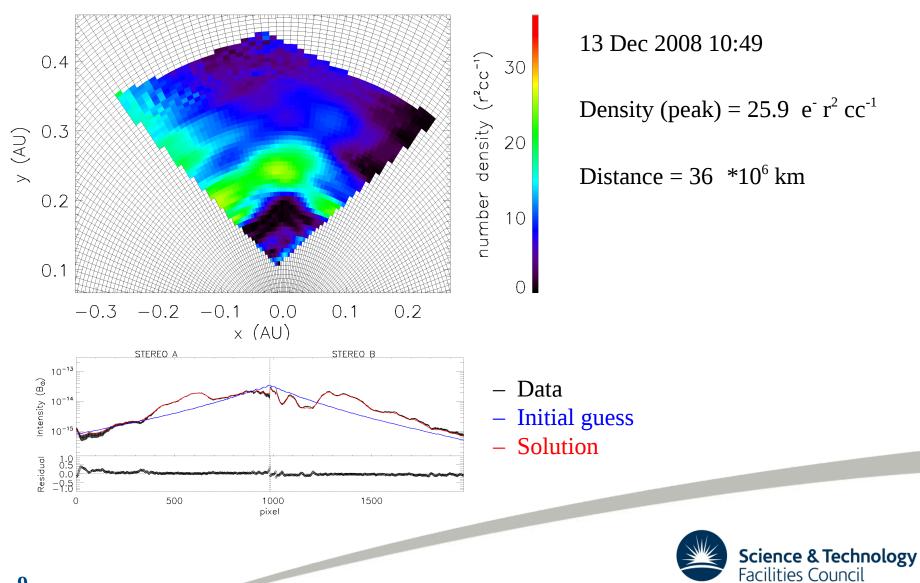


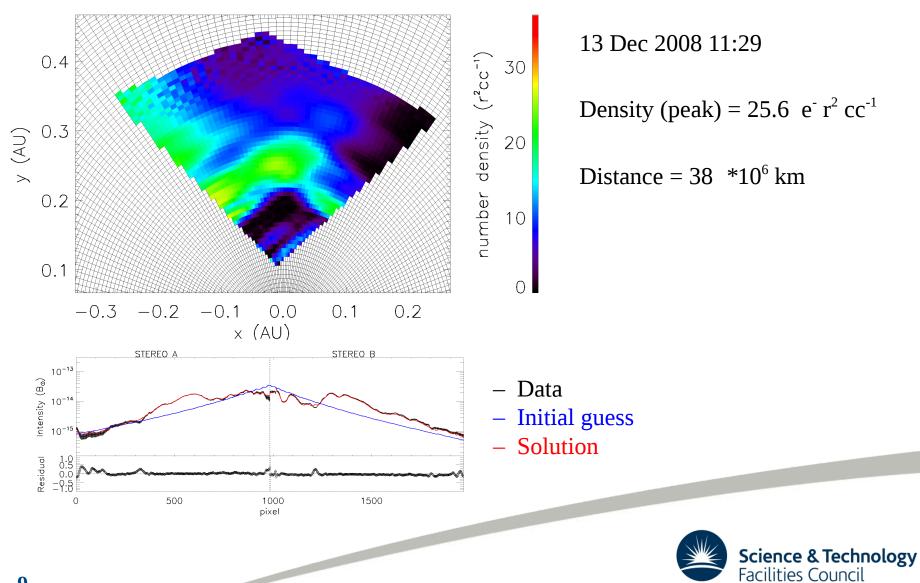


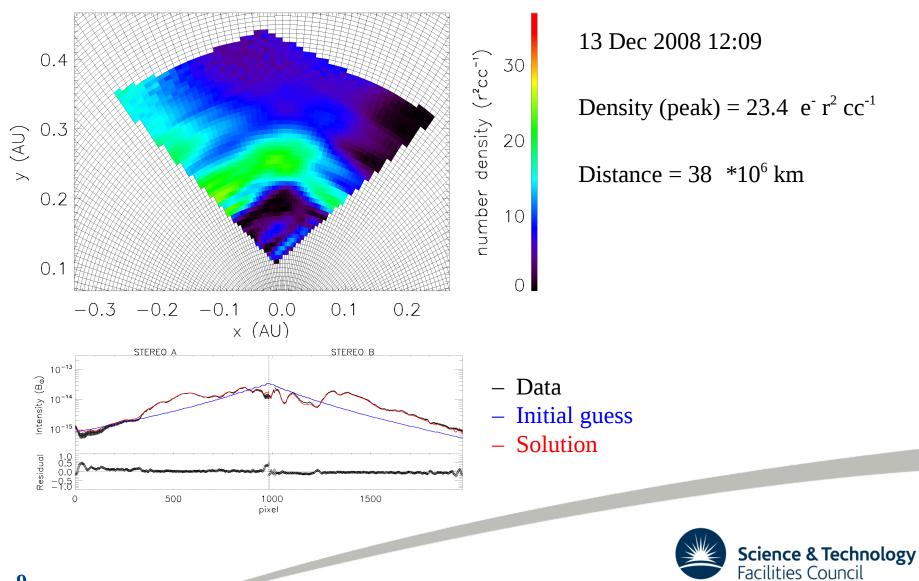


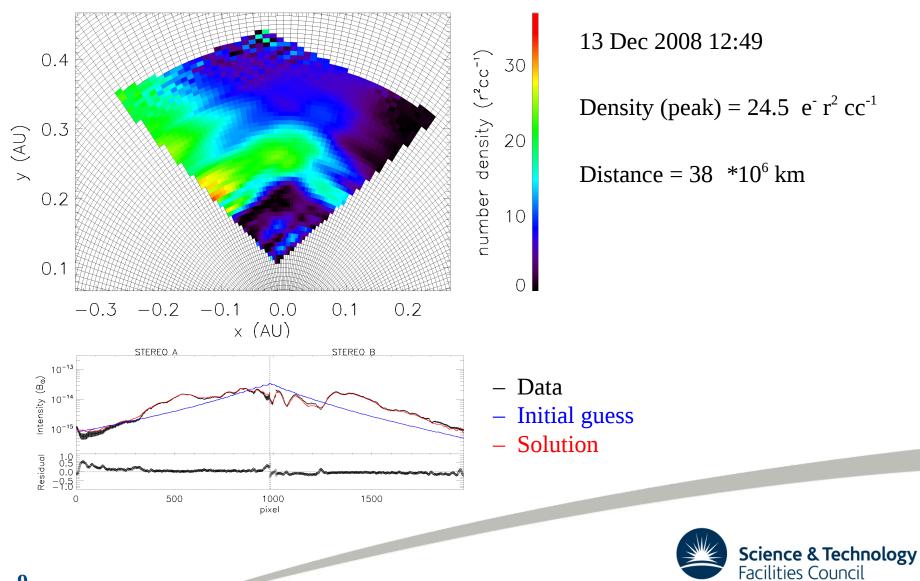


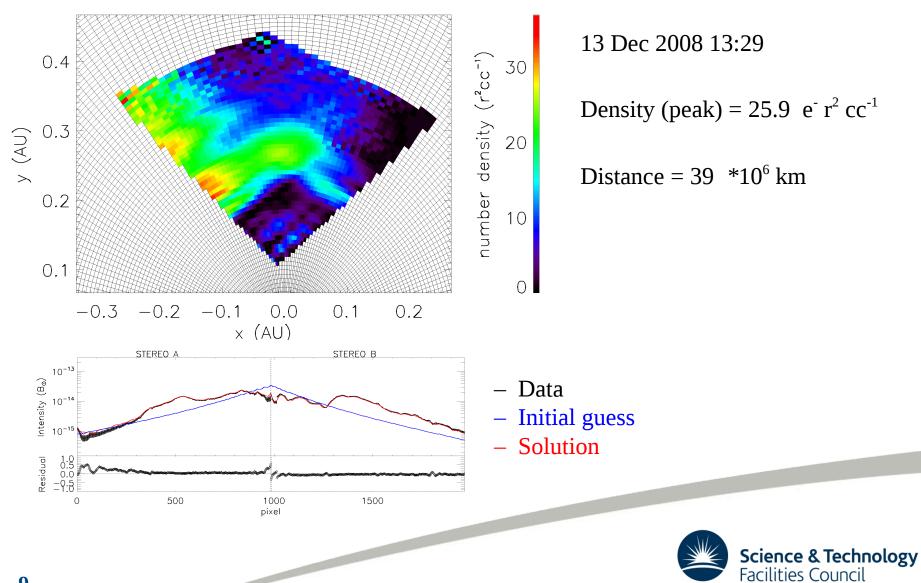


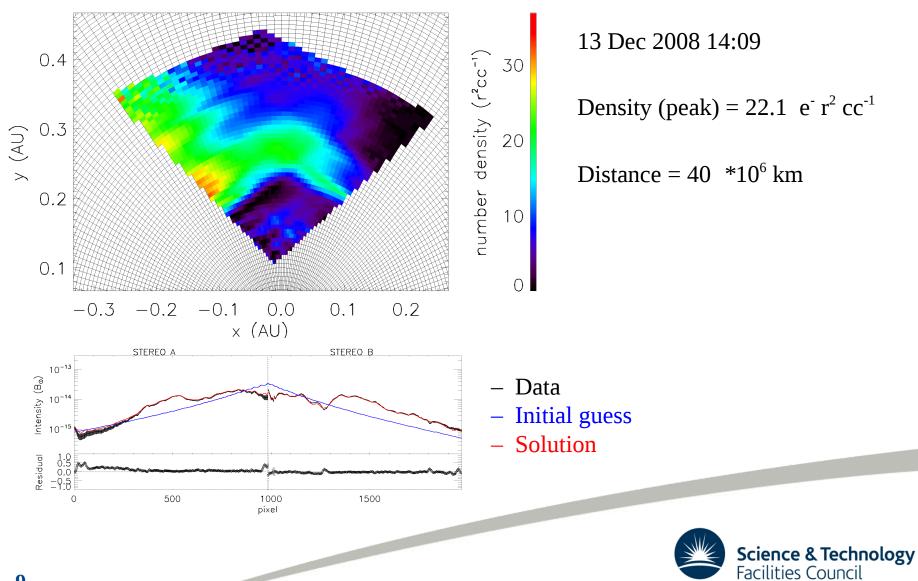


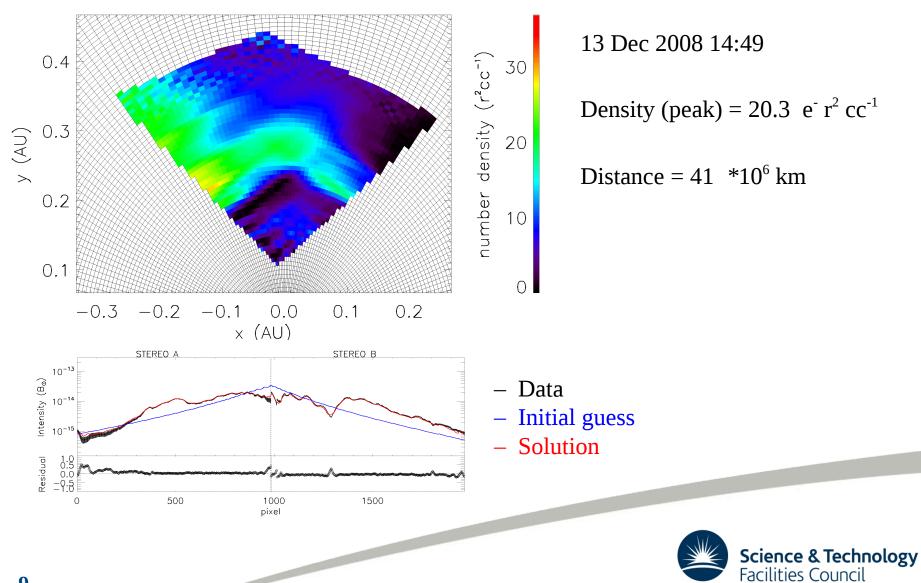


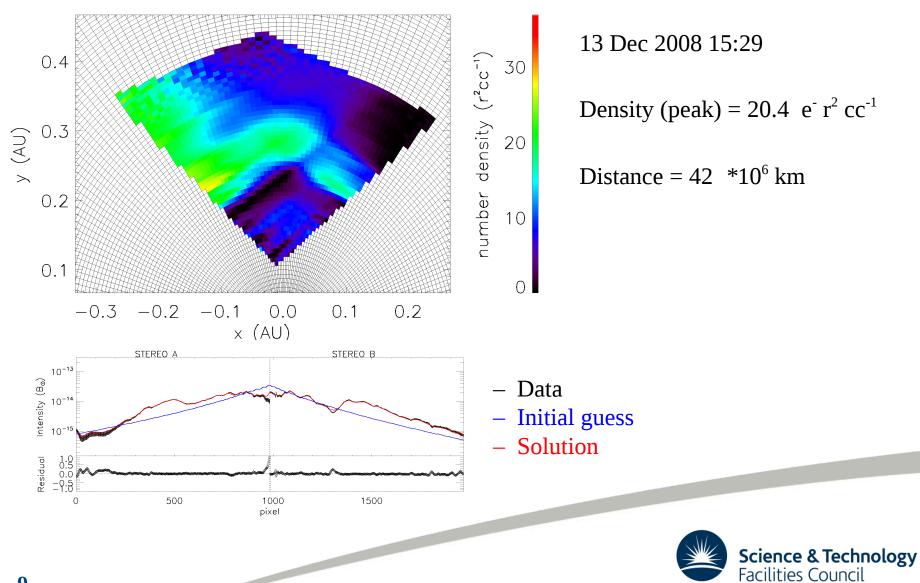


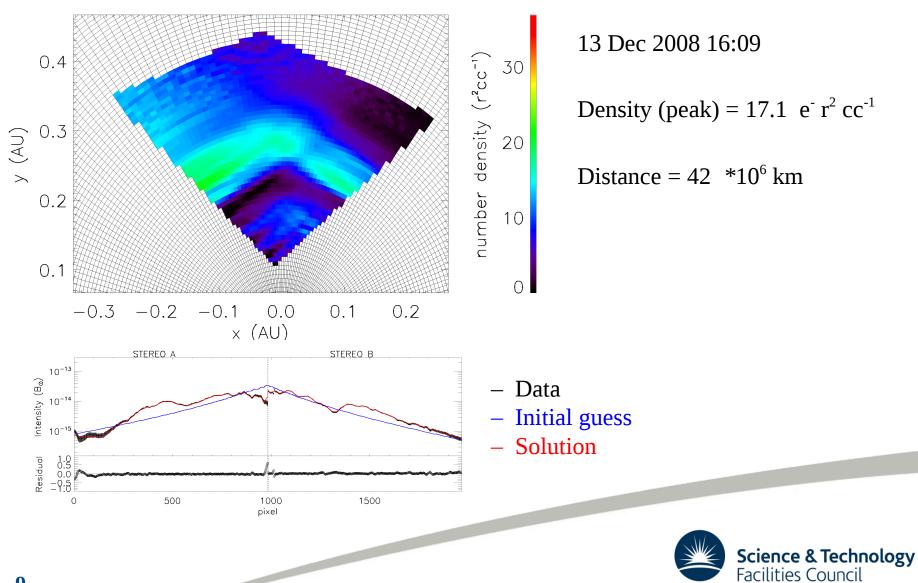


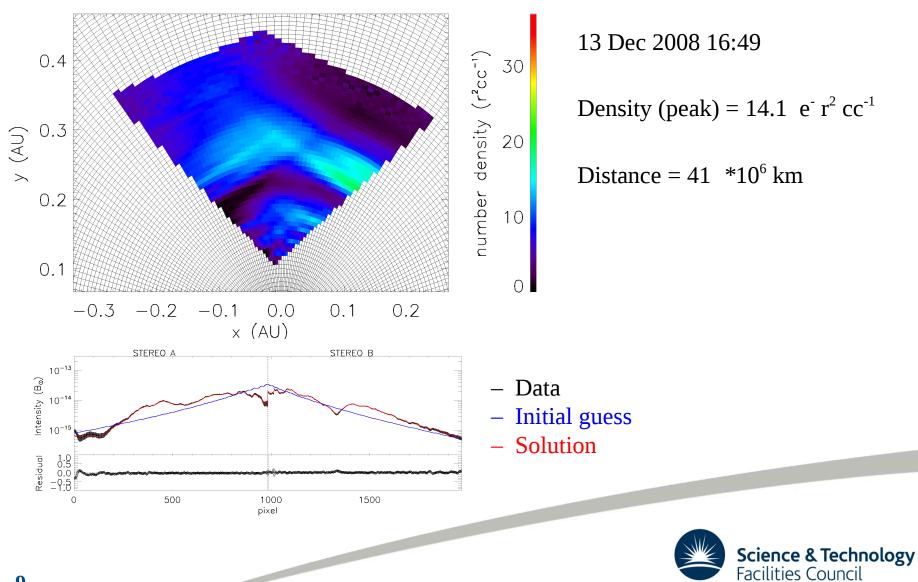


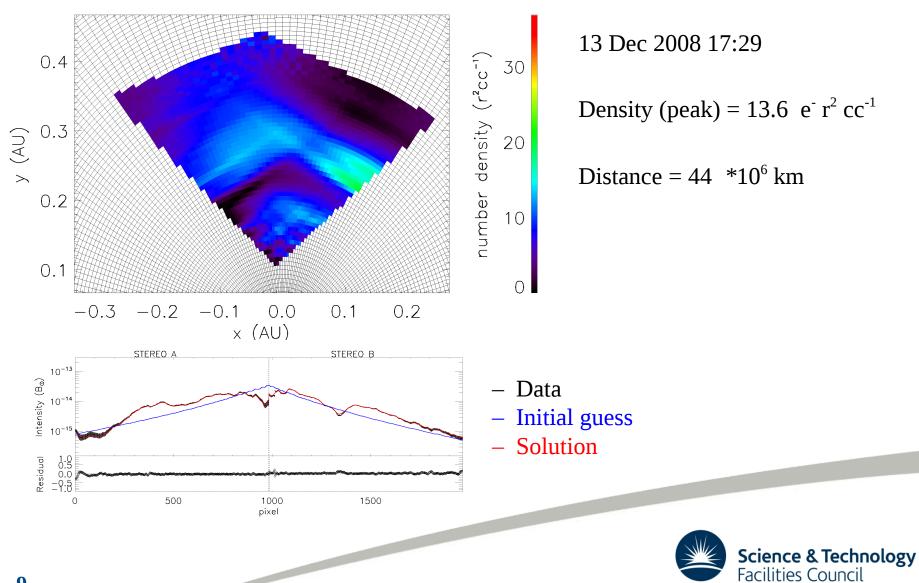


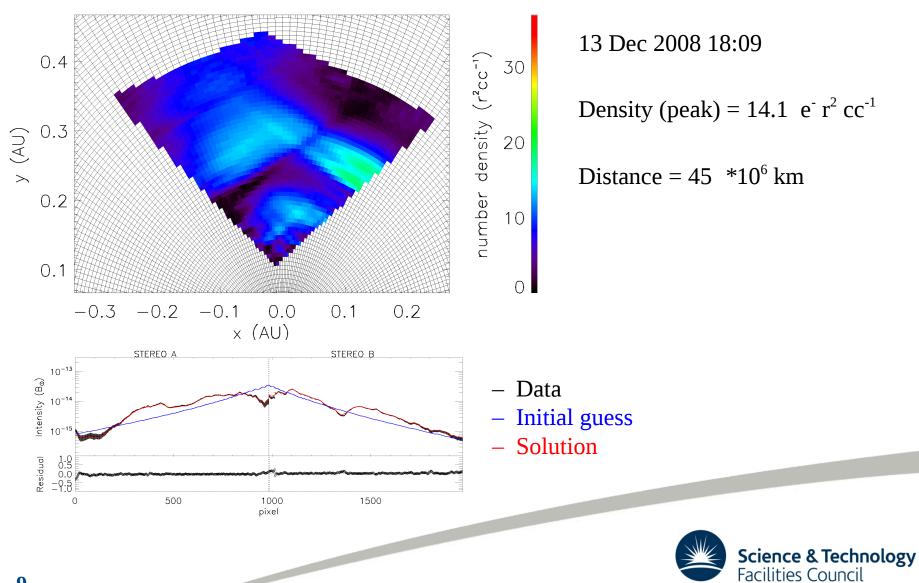


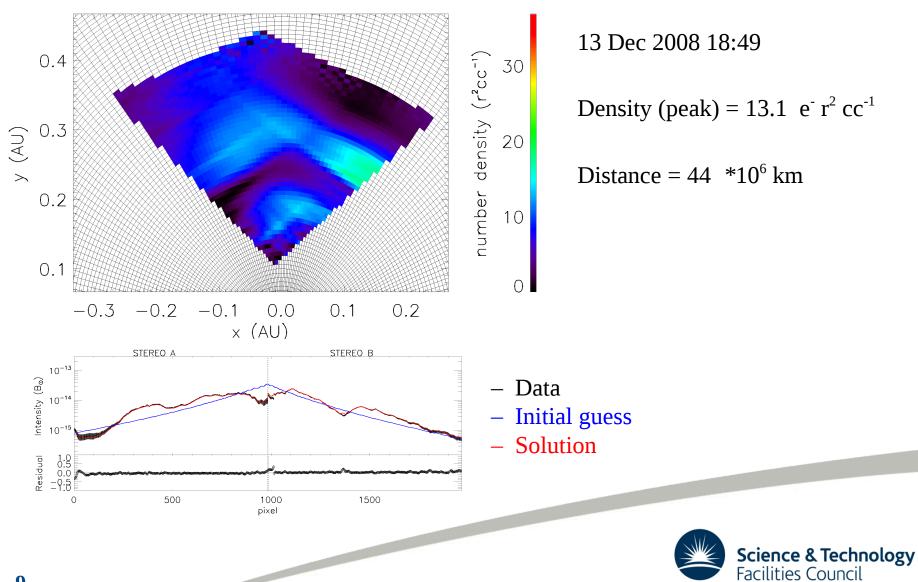




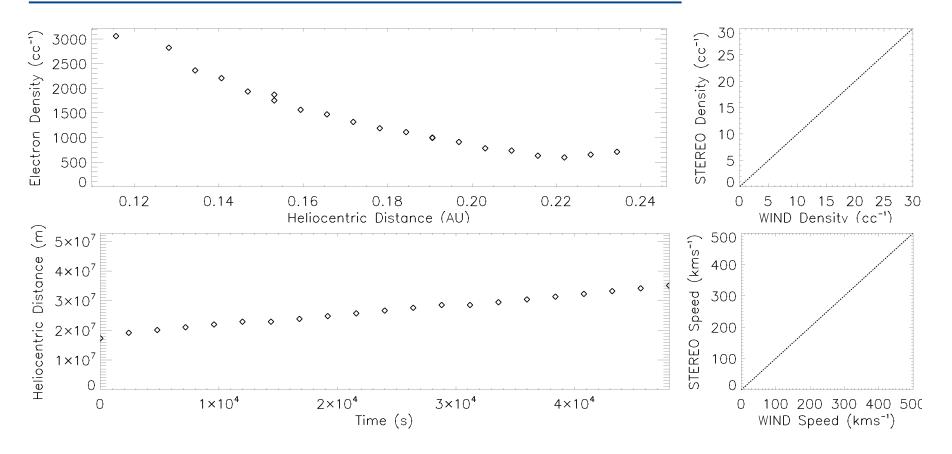






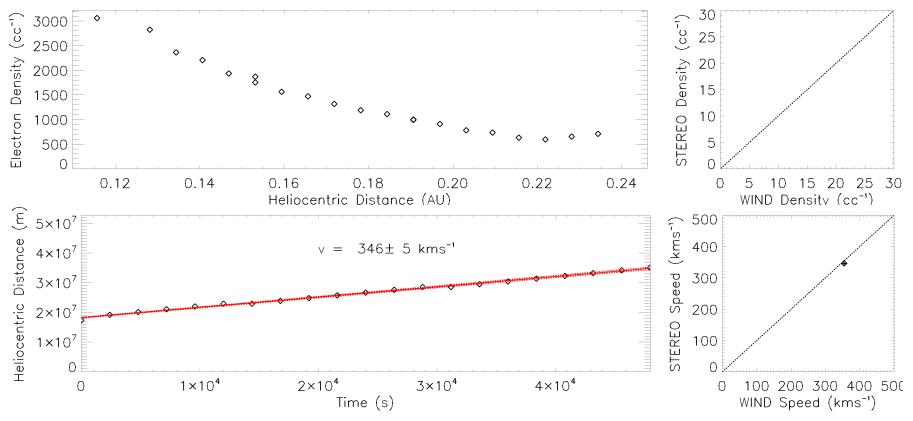


Dec 2008 CME Speed and Density





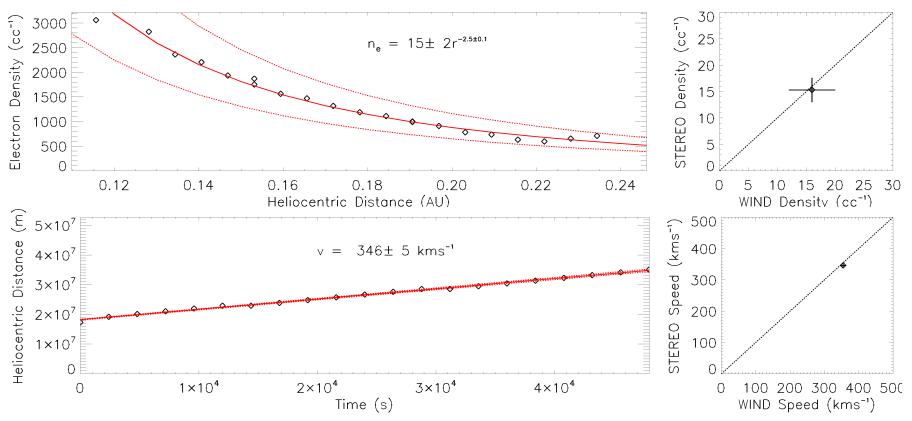
Dec 2008 CME Speed and Density



Assume constant velocity



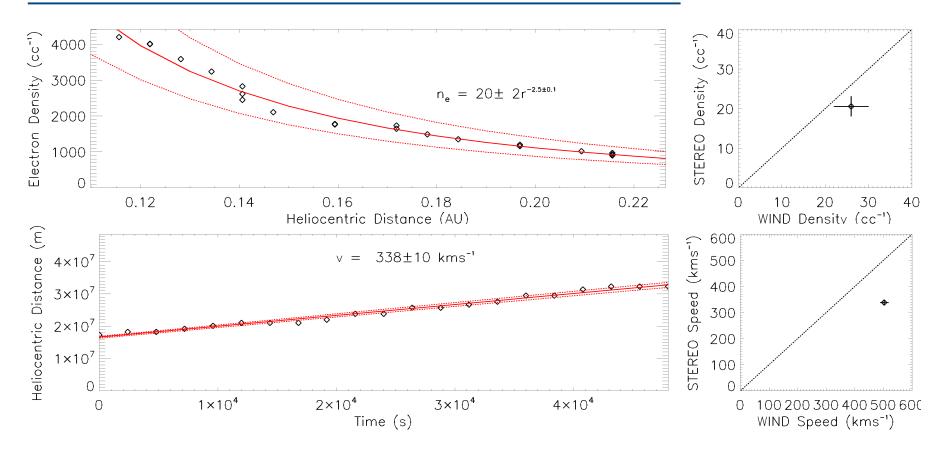
Dec 2008 CME Speed and Density



- Assume constant velocity
- Assume density and radius follow power law
- Regularisation terms can be tuned to give more accurate value for density

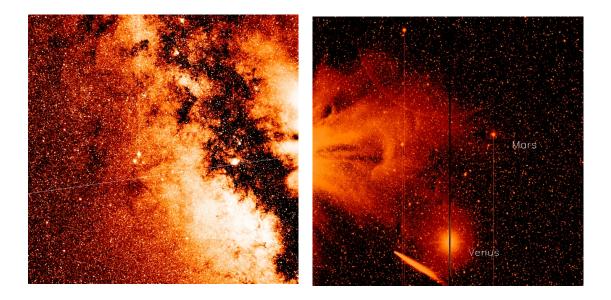


Oct 2011 CME Speed and Density





Limitations



- The galactic plane passes through the FOV of each spacecraft ~2 times per year
- Planets and comets frequently obscure images
- Relative position of the two spacecraft causes line-of-sight problems
- CME must be Earth-directed



Summary

- Electron density distribution in the ecliptic has been measured using combined HI data from both STEREO spacecraft
- The technique may be used if corrections are applied to produce a physically consistent K-corona intensity
- Iterative algorithm typically produces a density distribution with a mean residual of less than 5% of observations
- May be used to estimate CME speeds and densities
- Only two CMEs to which this technique has been applied

